

# Nuclear Radius Systematics around $N=Z$ Kr Isotopes Studied via their Interaction Cross-Sections at Relativistic Energies

T. Yamaguchi<sup>1</sup>, T. Suzuki<sup>1</sup>, T. Ohnishi<sup>2</sup>, K. Sümmerer<sup>3</sup>, F. Becker<sup>3</sup>, M. Fukuda<sup>4</sup>, H. Geissel<sup>3</sup>, M. Hosoi<sup>1</sup>, R. Janik<sup>5</sup>, A. Kelic<sup>3</sup>, K. Kimura<sup>6</sup>, S. Mandal<sup>3</sup>, G. Münzenberg<sup>3</sup>, S. Nakajima<sup>1</sup>, T. Ohtsubo<sup>7</sup>, A. Ozawa<sup>8</sup>, A. Prochazka<sup>5</sup>, M. Shindo<sup>9</sup>, B. Sitar<sup>5</sup>, P. Strmen<sup>5</sup>, T. Suda<sup>2</sup>, K. Sugawara<sup>1</sup>, I. Szarka<sup>5</sup>, A. Takisawa<sup>7</sup>, M. Takechi<sup>4</sup>, K. Tanaka<sup>2</sup>

<sup>1</sup> Department of Physics, Saitama University, Saitama City, Saitama 338-8570, Japan

<sup>2</sup> The Institute of Physical and Chemical Research (RIKEN), Wako-shi, Saitama 351-0198, Japan

<sup>3</sup> Gesellschaft für Schwerionenforschung (GSI), Planckstrasse 1, D-64291 Darmstadt, Germany

<sup>4</sup> Department of Physics, Osaka University, Toyonaka, Osaka 560-0043, Japan

<sup>5</sup> Faculty of Mathematics and Physics, Comenius University, 84215 Bratislava, Slovak Republic

<sup>6</sup> Nagasaki Institute of Applied Science, Nagasaki 851-0193, Japan

<sup>7</sup> Department of Physics, Niigata University, Niigata City, Niigata 950-2181, Japan

<sup>8</sup> Department Of Physics, University of Tsukuba, Tsukuba-shi, Ibaraki 305-8577, Japan

<sup>9</sup> Department of Physics, University of Tokyo, Bunkyo-ku, Tokyo 113-0033, Japan

Precision measurements of the interaction cross-sections ( $\sigma_I$ ) at relativistic energies  $\sim 1$  AGeV allow us to derive nuclear matter radii [1]. Since nuclear matter radii are directly related to the density distributions, the measurements of  $\sigma_I$  are a good probe to search for unusual nuclear structures. Neutron-deficient Kr isotopes close to the  $N = Z$  line attract a particular interest in their structures. Recent  $\gamma$ -ray measurements have shown up that large deformations and even shape coexistence appear in this region [2]. The proton dripline nuclei in the mass range  $60 < A < 100$  play an important role to determine the astrophysical rapid proton capture process path [3].

We have performed the precision measurements of the interaction cross-sections for neutron-deficient even-mass Kr isotopes using the fragment separator FRS at GSI. We have so far succeeded several experiments to determine nuclear matter radii of light nuclei ( $A \leq 40$ ) [4-6]. The measurements for Kr isotopes are an effort to extend a series of successful experiments to heavier nuclei ( $A \sim 80$ ). To identify the particles unambiguously we have developed the ultra-fast timing plastic scintillation counters with the timing resolution of 13 ps ( $\sigma$ ) for  $^{80}\text{Kr}$  beam at 1.05 AGeV.

The result of the interaction cross-sections of Kr isotopes clearly shows that the radii increase as the mass number decreases. This is clearly different from our early observations in the light region [6]. A systematics of the radius changes in the medium mass region around the  $N = Z$  line is discussed together with the results of indirect measurements recently performed at GANIL [7]. Since the charge radii of Kr isotopes are available from the isotope-shift measurements at ISOLDE [8], the proton and the neutron radius can be extracted separately. The proton skin thickness for the neutron-deficient Kr isotopes is for the first time compared with that observed in the light nuclei.

The present experiment extended the database of the production cross-sections for exotic nuclei by studying the fragmentation of  $^{80}\text{Kr}$ . A systematics of the production cross-sections and the charge pickup channels is compared with the reaction models such as abrasion-ablation model [9].

[1] I. Tanihata *et al.*, Phys. Lett. B 287 (1992) 307.

[2] A. Gade *et al.*, Phys. Rev. Lett. 95 (2005) 022502; E. Bouchez *et al.*, Phys. Rev. Lett. 90 (2003) 082502.

[3] B. Blank *et al.*, Phys. Rev. Lett. 74 (1995) 4611; M.J. López Jiménez *et al.*, Phys. Rev. C 66 (2002) 025803.

[4] T. Suzuki *et al.*, Phys. Rev. Lett. 75 (1995) 3241; Nucl. Phys. A 658 (1999) 313.

[5] L.V. Chulkov *et al.*, Nucl. Phys. A 674 (2000) 330; Nucl. Phys. A 603 (1996) 219.

[6] A. Ozawa *et al.*, Nucl. Phys. A 691 (2001) 599; Nucl. Phys. A 709 (2002) 60.

[7] G.F. Lima *et al.*, Nucl. Phys. A 735 (2004) 303.

[8] M. Keim *et al.*, Nucl. Phys. A 586 (1995) 219.

[9] J.J. Gaimard and K.-H. Schmidt, Nucl. Phys. A 531 (1991) 709.