

# The lowest excited states in $^{13}\text{O}$

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Spectroscopy of exotic nuclei far from stability is crucial to understanding of nuclear structure. The exotic nucleus,  $^{13}\text{O}$ , is also of potential importance for the p-p chain of reactions in population III metal poor stars via the  $(p,\gamma)$  reaction on  $^{12}\text{N}$  [1]. We have measured the lowest excited states in  $^{13}\text{O}$  for the first time using the inverse kinematics thick target method [2]. The experiment was carried out on TwinSol radioactive beam facilities [3] at the University of Notre Dame nuclear structure laboratory. Elastic proton scattering on  $^{12}\text{N}$  was measured in the laboratory energy range from 4 - 10 MeV and analyzed in the framework of the **R**-matrix formalism and with a potential model. The two lowest excited states in  $^{13}\text{O}$  have the spin and parities of  $1/2^+$  at 2.69 MeV and  $(1/2^-, 3/2^-)$  at 3.29 MeV. Single particle spectroscopic factors were defined for these states based on different assumptions for a shell model potential of  $^{13}\text{O}$ . On the basis of the obtained data, we calculate the relevant astrophysical reaction rate of  $^{12}\text{N}(p,\gamma)^{13}\text{O}$  process and discuss its implications for population III stars.

[1] M. Wiescher et al., *Astro. J.*, 343, 352 (1989);

[2] K. P. Artemov et al., *Sov. Nucl. Phys.*, 52, 406, (1990);

[3] M.Y. Lee et. al., *Nucl. Instrum. Meth. A*, 422, 536, (1995);